

An Alternative Formulation of Rigorous Mean-Field Theory

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It is shown how to derive rigorous mean-field theory from a type of many-body interaction.

KEY WORDS: Mean-field theory; many-body interactions.

Since the work of Kac *et al.*⁽¹⁾ in 1963, the classical or mean-field theories of a variety of systems have been derived rigorously. They include the three-dimensional version and extension^(2,3) of Ref. 1, quantum mechanical extensions,^(4,5) antiferromagnet and melting models,⁽⁶⁾ classical correlation functions,⁽⁷⁾ and theories of amorphous ferromagnets⁽⁸⁾ and metastable states.⁽⁹⁾

The derivations are valuable because of the extensive use of mean-field theories in the literature. One objection to the theories is their inapplicability to the critical region of a phase transition. In particular they yield inaccurate values for critical exponents. On a more theoretical level they lack some appeal (see the introductory comments in Ref. 11) because the transitions are in a sense "forced" by the application of a long-range limit operation on the thermodynamic functions. We outline here how this latter problem may be formally removed by an alternative formulation in terms of a type of many-body interaction.

We take the total potential energy of N particles at points $\mathbf{x}_1, \mathbf{x}_2, \dots, \mathbf{x}_N$ to be the sum of two-body, three-body, up to N -body interactions:

$$U_N(\mathbf{x}_1, \dots, \mathbf{x}_N) = \sum_{1 \leq i < j \leq N} [q(\mathbf{x}_i - \mathbf{x}_j) + \Phi_2(\mathbf{x}_i - \mathbf{x}_j)] + \sum_{1 \leq i < j < k \leq N} \Phi_3(\mathbf{x}_i, \mathbf{x}_j, \mathbf{x}_k) + \dots + \Phi_N(\mathbf{x}_1, \dots, \mathbf{x}_N) \quad (1)$$

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where

$$\Phi_m(\mathbf{x}_1, \dots, \mathbf{x}_m) = \sum_{1 \leq i < j \leq m} \psi_m(\mathbf{x}_i - \mathbf{x}_j) \quad (2)$$

$$\psi_m(\mathbf{r}) = \sum_{k=2}^m (-1)^{m-k} \binom{m-2}{k-2} \gamma_k^\nu \varphi(\gamma_k \mathbf{r}) \quad (3)$$

and

$$\gamma_k = k^{-\eta/\nu} \quad (4)$$

where ν is the dimensionality of the system and η is a constant satisfying

$$0 < \eta < 1 \quad (5)$$

Our main result is that if $q(\mathbf{r})$ and $\varphi(\mathbf{r})$ satisfy the conditions imposed by Lebowitz and Penrose, then the pressure in the thermodynamic limit is given by the Maxwell construction or equal area rule⁽²⁾ applied to the function

$$p^0(\rho, T) + \frac{1}{2}\alpha\rho^2 \quad (6)$$

where ρ is the density, T is the temperature, p^0 is the pressure of a system with two-body potential $q(\mathbf{r})$ alone, and

$$\alpha = \int d\mathbf{r} \varphi(\mathbf{r}) \quad (7)$$

the integral being over all of ν -dimensional space. The variational principle of Penrose and Gates holds under the weaker conditions on φ given in Ref. 3.

To establish these results, we substitute (2) and (3) in (1) and obtain

$$U_N(\mathbf{x}_1, \dots, \mathbf{x}_N) = \sum_{1 \leq i < j \leq N} [q(\mathbf{x}_i - \mathbf{x}_j) + \gamma_N^\nu \varphi\{\gamma_N(\mathbf{x}_i - \mathbf{x}_j)\}] \quad (8)$$

We therefore have a genuine two-body interaction through $q(\mathbf{r})$ and an artificial two-body potential $\gamma_N^\nu \varphi(\gamma_N \mathbf{r})$ which depends on the total number N of particles in the system. As N increases, the range γ_N^{-1} of interaction also increases.

This may be compared with the model of Lebowitz and Penrose,⁽²⁾ who use a two-body potential

$$q(\mathbf{r}) + \gamma^\nu \varphi(\gamma \mathbf{r})$$

where γ is an arbitrary positive parameter. They then take the long-range limit $\gamma \rightarrow 0$ after the thermodynamic limit $N \rightarrow \infty$.

We note from (4) that

$$\begin{aligned} \gamma_N &\rightarrow 0 & \text{as } N &\rightarrow \infty \\ N\gamma_N^\nu &\rightarrow \infty & \text{as } N &\rightarrow \infty \end{aligned} \quad (9)$$

Consequently, we have

$$r_0 \ll \gamma_N^{-1} \ll N^{1/\nu} \quad (10)$$

for large N , where r_0 is the hard-core diameter of $q(\mathbf{r})$. Hence we can divide the container into cells of volume ω satisfying

$$r_0 \ll \omega \ll \gamma_N^{-1} \ll N^{1/\nu} \quad (11)$$

for large N . The inequalities of Lebowitz and Penrose therefore apply with γ_N replacing their parameter γ . The result (6) then follows by their method.

Our formulation gives (6), and the consequent van der Waals–Maxwell phase transition, without the use of the additional limiting operation $\gamma \rightarrow 0$. The present work is related to the extensive work of Fisher⁽¹⁰⁾ and Fisher and Felderhof,⁽¹¹⁾ which describes the phase transitions in a class of one-dimensional models with many-body interactions. However, our model does not quite belong to the class considered by these authors, because our potentials Φ_m do not satisfy their tempering conditions.

On the other hand, our results provide an example of a system which violates the usual conditions on interactions, but for which the thermodynamic limit exists and the pressure is a nondecreasing, continuous function of the density ρ . Thus the usual conditions are sufficient, but not necessary.

The same formulation is applicable to other mean-field models. It may be contrasted with the Temperly model, which is a lattice gas (or Ising model) with interaction energy

$$(1/V) \sum_{1 \leq i < j \leq V} \sigma_i \sigma_j \quad \text{and} \quad \sum_{i=1}^V \sigma_i = N \quad (12)$$

where $\sigma_i = 0$ or 1. This yields the van der Waals theory *without* the Maxwell construction. As the number of lattice sites V tends to infinity, the interaction decreases at the rate $1/V$, while the thermodynamic limit imposes the condition $N/V \rightarrow \text{constant}$, ρ . Consequently it is not possible to divide the system into cells in such a way that the interaction in a single cell is negligible, while each cell contains an arbitrarily large number of particles in the thermodynamic limit. Each cell contributes a factor $\frac{1}{2}\rho^2$. In the new model (and that of Lebowitz and Penrose) the factor $\frac{1}{2}\alpha\rho^2$ in (6) results from interactions between such cells. It is this feature, expressed through (11), which yields the Maxwell construction.

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